Current Progress of 380 cm Telescope and Future Research Plan in Timau National Observatory

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NATIONAL RESEARCH AND INNOVATION AGENCY (BRIN)

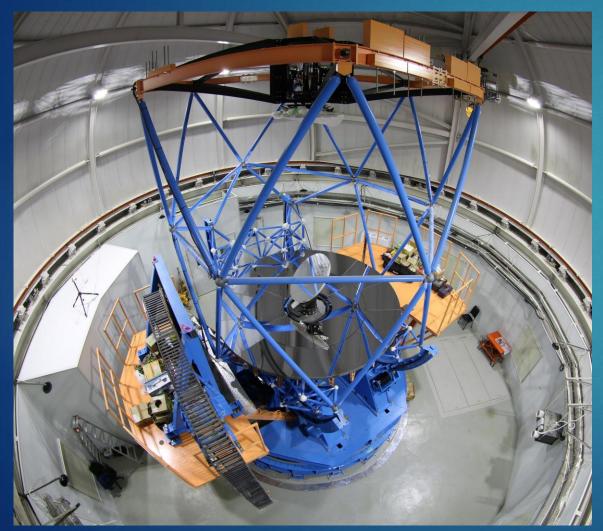
INDONESIA

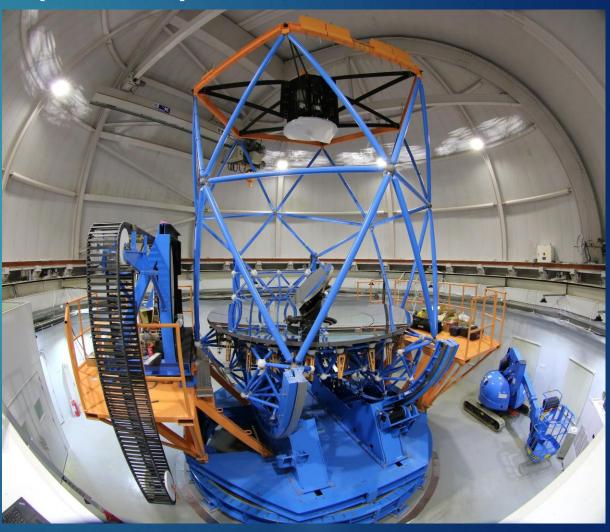














North camera



Dome camera down



East camera



South camera



Current Status:

- All mirrors have been installed (M1 and M3 in March 2025).
- Mirror setting will be done in the end of September 2025.
- Optika will be installed in the end of September 2025
 Engineering First Light.
- > Scientific First Light is expected by the end of 2025.
- NIRCa will be installed in 2026.
- Research projects are expected to be started in 2026B (domestic and collaboration partners).

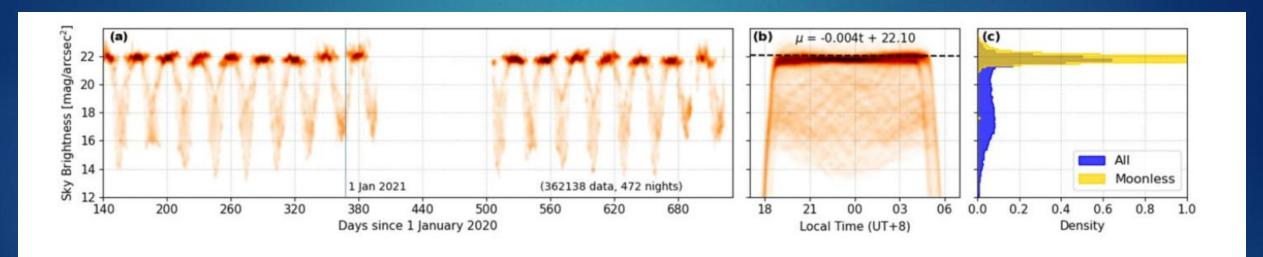


Figure 2. Density plot of the sky brightness over Timau as a function of the number of days since 2020 January 1 (panel a) and as a function of the local time (panel b). One-dimensional distribution of the sky brightness is also presented (panel c).

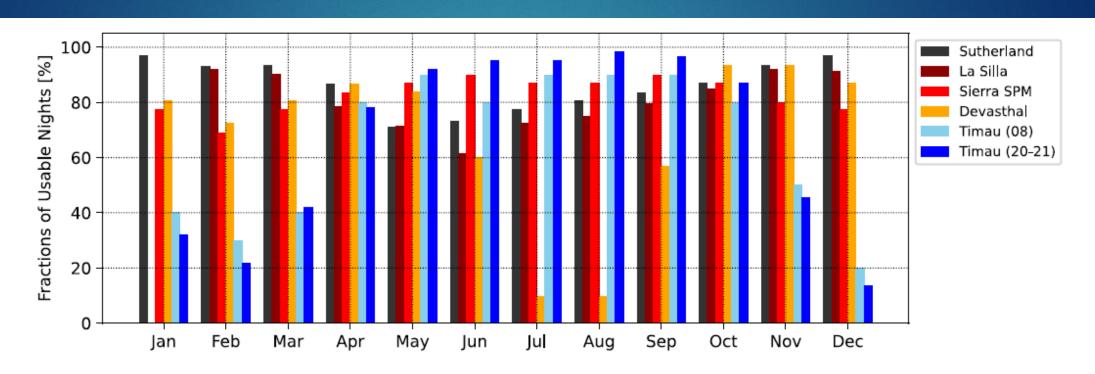


Figure 10. The percentages of usable nights at Timau and some other astronomical sites. Timau (20–21) is from the current study, Timau (08) data is from Hidayat et al. (2012), La Silla is from Cavazzani et al. (2020), while the Sutherland, Devasthal, and Sierra San Pedro Martir are from ERA5 reanalysis by Ningombam et al. (2021).

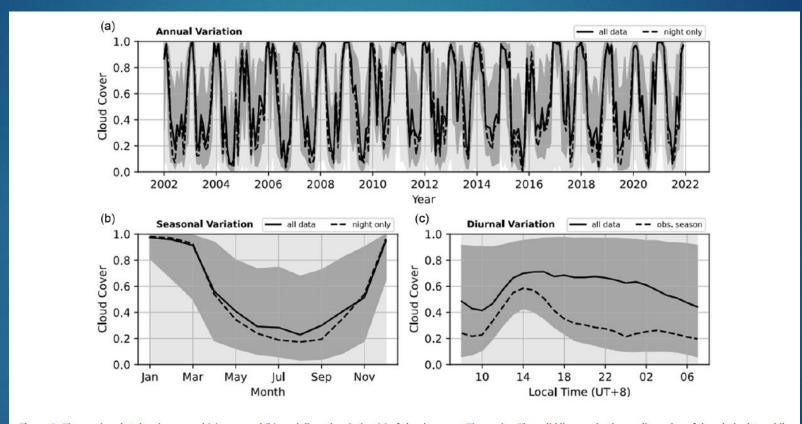
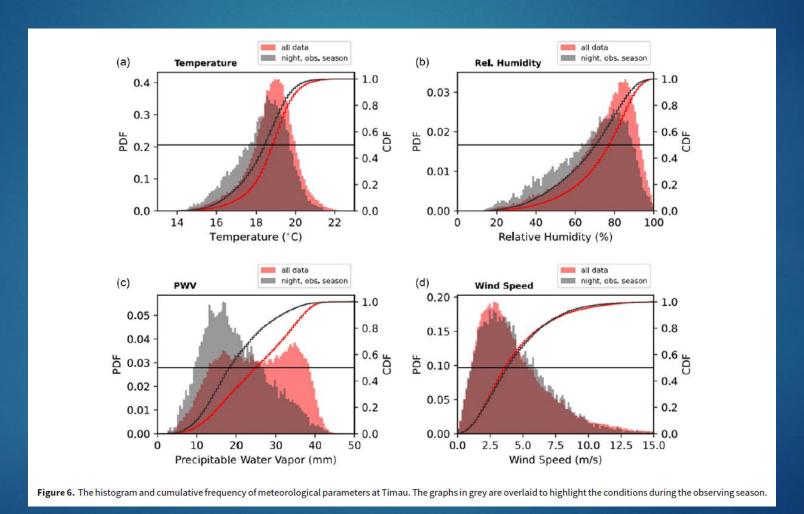
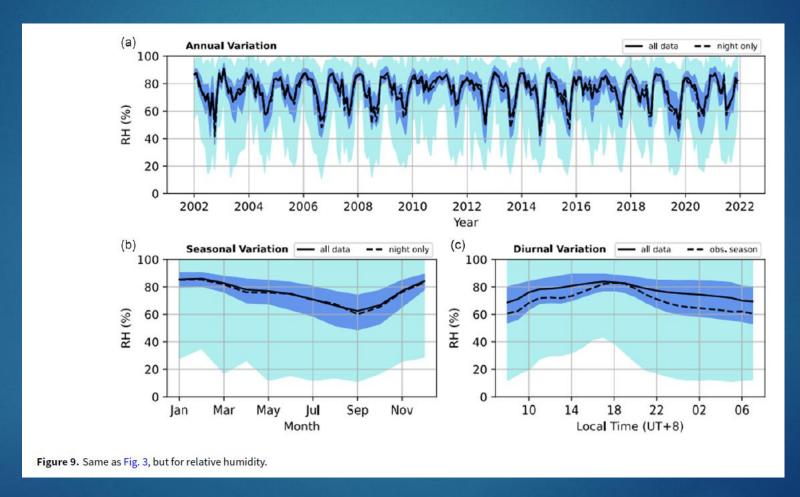


Figure 3. Time series plot showing annual (a), seasonal (b), and diurnal variation (c) of cloud cover at Timau site. The solid line marks the median value of the whole data while the dashed line is for the time-filtered data. The shaded area encompasses the interquartile range (percentile 25–75) of the temperature whereas the minimum to maximum range is lightly shaded.





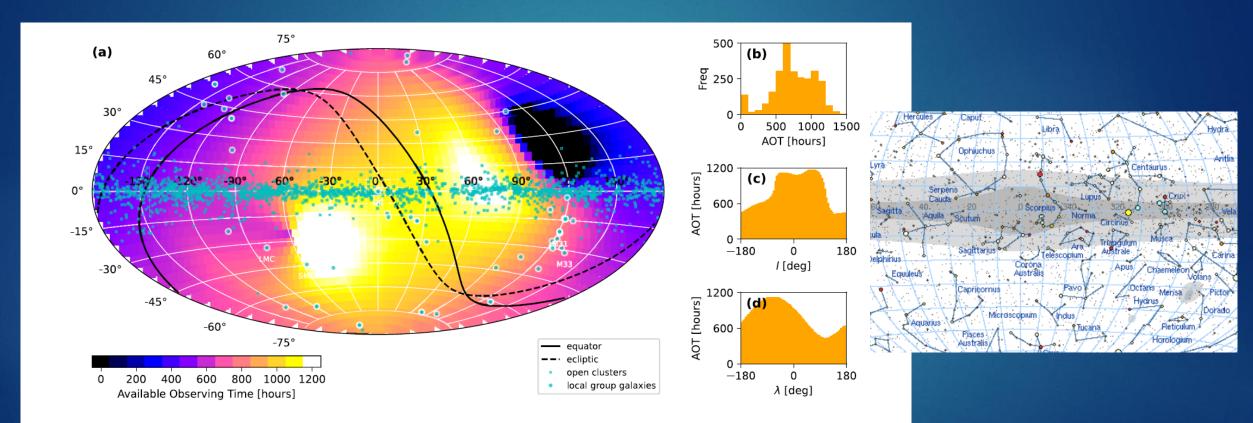


Figure 11. The observability map for the Timau shows the AOT of sky segments in Galactic coordinate (panel a). The location of open clusters (Dias et al. 2021) and local group galaxies are indicated by crosses and circles. The number distribution of sky segments as a function of AOT is also provided (panel b). The average AOT of sky segments as a function of galactic longitude (l) and ecliptic longitude (λ) are provided in panels (c) and (d). The average values are from the segments within $\pm 30^{\circ}$ from the galactic equator or ecliptic.

Timau telescope ~ Seimei telescope

- Two Cameras: 3Optika dan NIRKA.
 - 3Optika for imaging and photometry has 3 bands:

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g (450-530 nm), r (550-680 nm), and i (700-820 nm).
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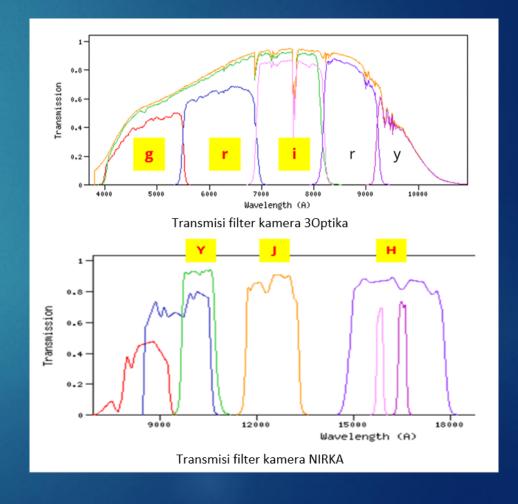
FoV: 12' x 12'.

"NIRCA" (Near IR Camera) has 3 bands:
 Y (~1020 nm), J (~ 1220 nm), and

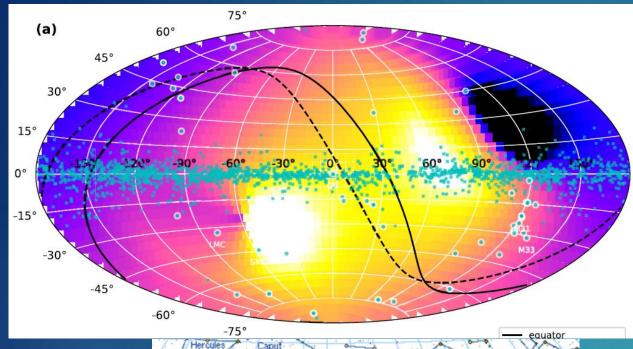
 $H(\sim 1530 \text{ nm}).$

FoV: 8.74' x 8.74'.





Future Research at TNOI



Hercules Caput Libra Hydra

Ophiuchus Libra Centaurus

Serpens Cauda 20 D Scorpius 340 S20 Ophiu Norma Circinus

sula Sagittarius Telescopium Australe Carina

Delphinus Capricornus Pavo Octaris Merisa Pictor

Hydra Dorado

Microscopium Indus Reticulum Musca Dorado

Apus Chaemeleon Volans

Capricornus Pavo Octaris Merisa Pictor

Hydrus Dorado

Apus Pisces Australis Horologium

Microscopium Indus Reticulum

Apus Pisces Australis

- Stellar astronomy:
 - Supernova
 - Late type stars flare
- Galactic astronomy
 - Star forming regions near the Galaxy Center
- Extra-galactic astronomy
- Planetary astronomy
 - Extrasolar planets
 - Asteroids
 - Comets