

Tricolor simultaneous photometric observations of tiny Near-Earth Objects (NEOs) with Seimei/TriCCS modified

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Seimei telescope

@ Okayama Observatory, Kyoto University

Overview

- We conducted tricolor simultaneous photometric observations.

- ▶ Instrument : Seimei telescope / TriCCS camera (g + r + i or z)
- ▶ Mode : video observation at 1 fps
- ▶ Target : 5 tiny Near-Earth Objects (diameter < 100 m)

- Rotational periods and spectral types were derived.

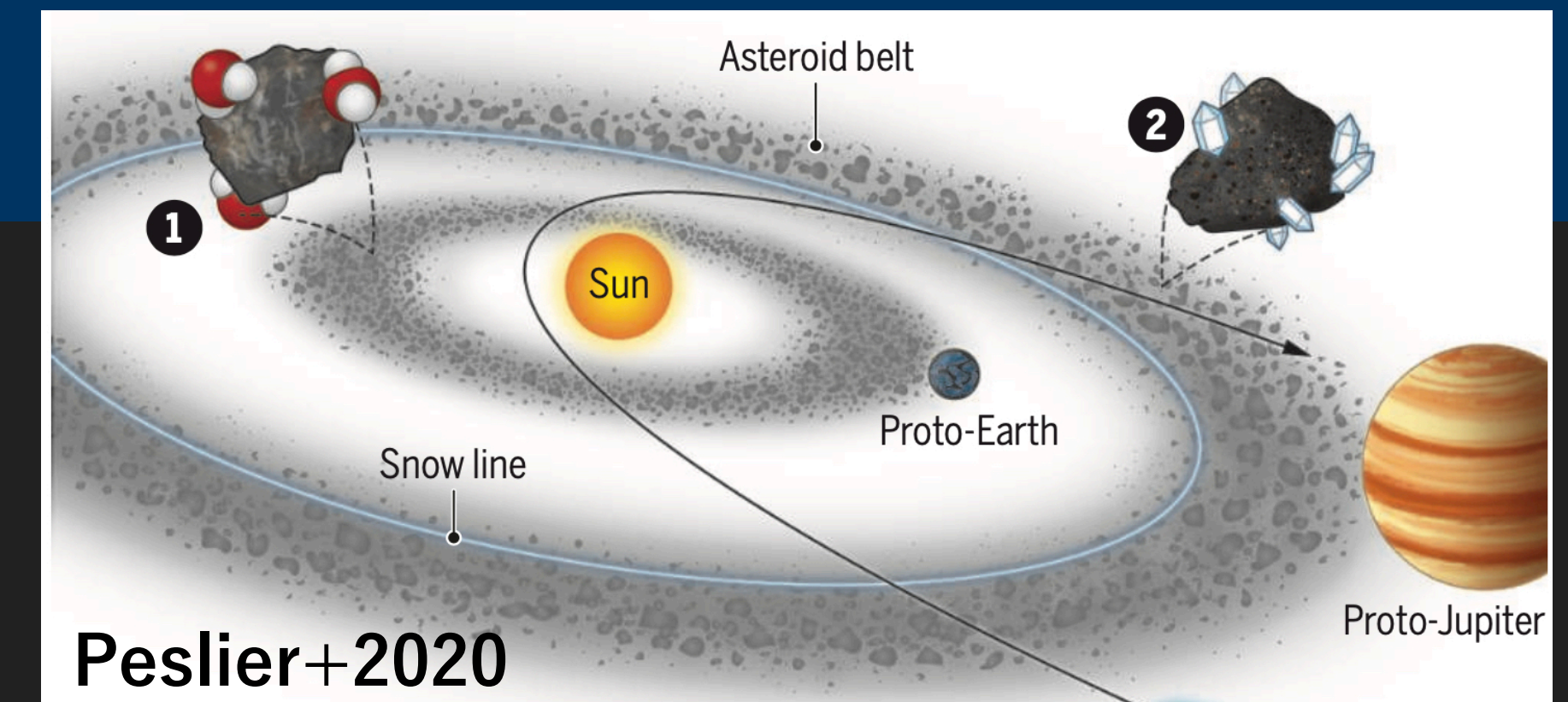
- ▶ Light curve -> rotational period
(+ light curve of a reference star)
Color -> spectral type
- ▶ We fully characterized 1 NEO.
- ▶ No spectral change was found in all 5 NEOs' light curves.



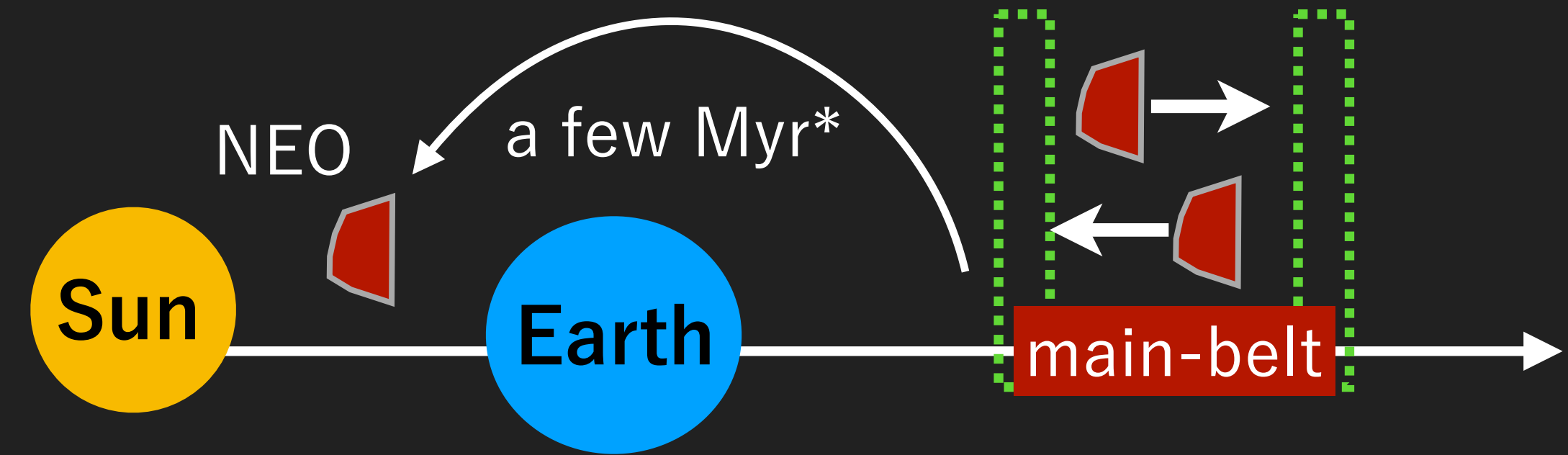
- Additional observations will be done in Semester 2021B.

Scientific motivation

- Material transportation by asteroids is related to the origin of the water and/or the life on the Earth.



- Near-Earth Objects** (NEOs, perihelion distance $q < 1.3$ au) come from the main-belt (~ 10 Myr).
- NEOs can deliver materials to the Earth.



- Spectroscopic or **multi-band photometric observations** reveal the surface composition.

- ▶ **Stony** : **S-type**
- ▶ **Carbonaceous** : **C-type**

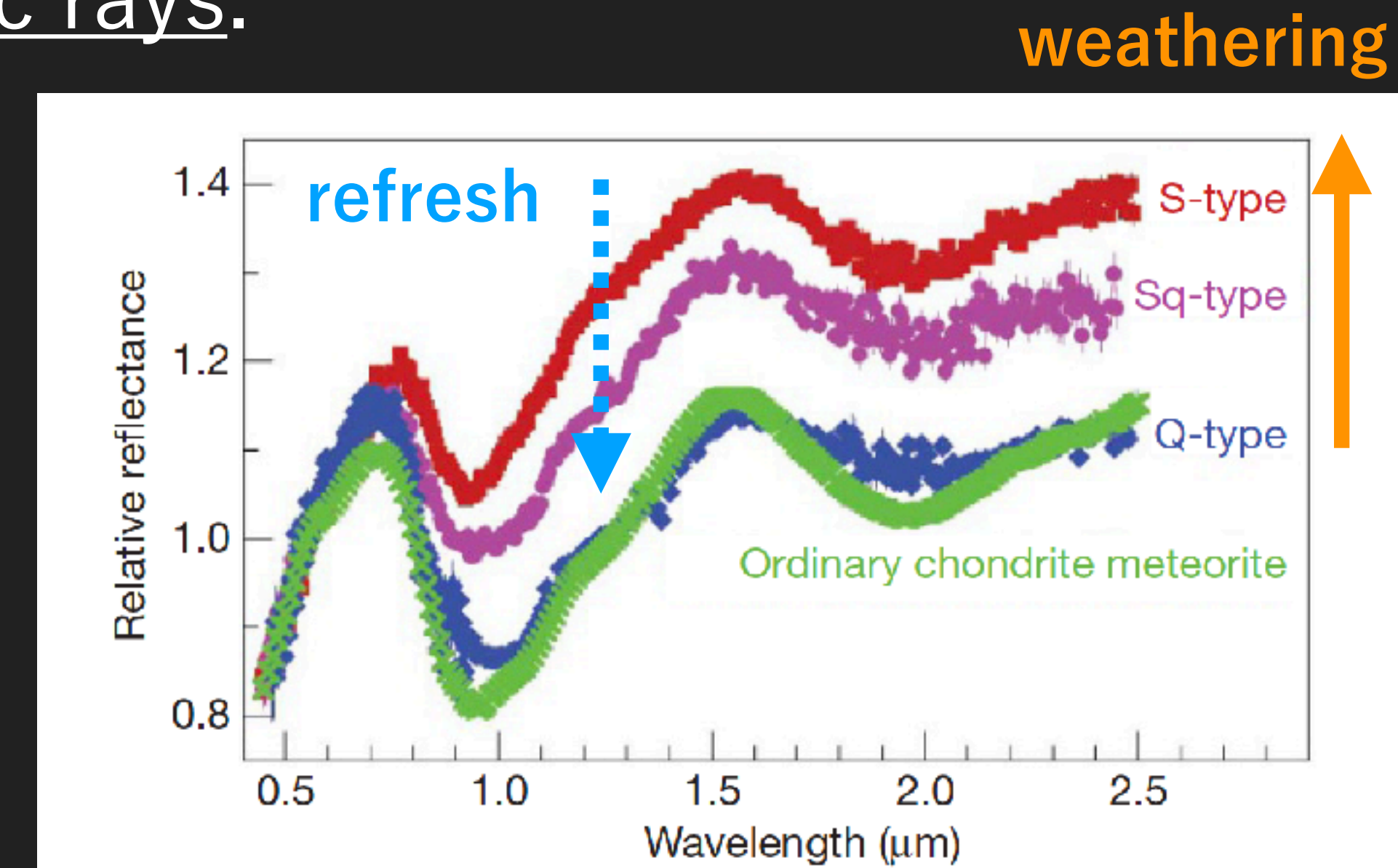


reflectance (0.45 – 2.45 micron)

From DeMeo+2009, Icarus, 202, 160, Fig.15

Space weathering and surface refreshing

- The surface property is altered by the **space weathering** : micro-meteorite impacts, solar wind, and cosmic rays.
- Asteroids' surfaces can be **refreshed** by other mechanisms, such as close encounters with planets and rotational fission.



Binzel+2010, Nature let., 463, 331, Figure 1

- **Both space weathering and its refresh are being discussed.**
 - ▶ Which weathering agents are dominant : micro-meteorite impact vs. solar wind?
 - ▶ Is there any size dependence?

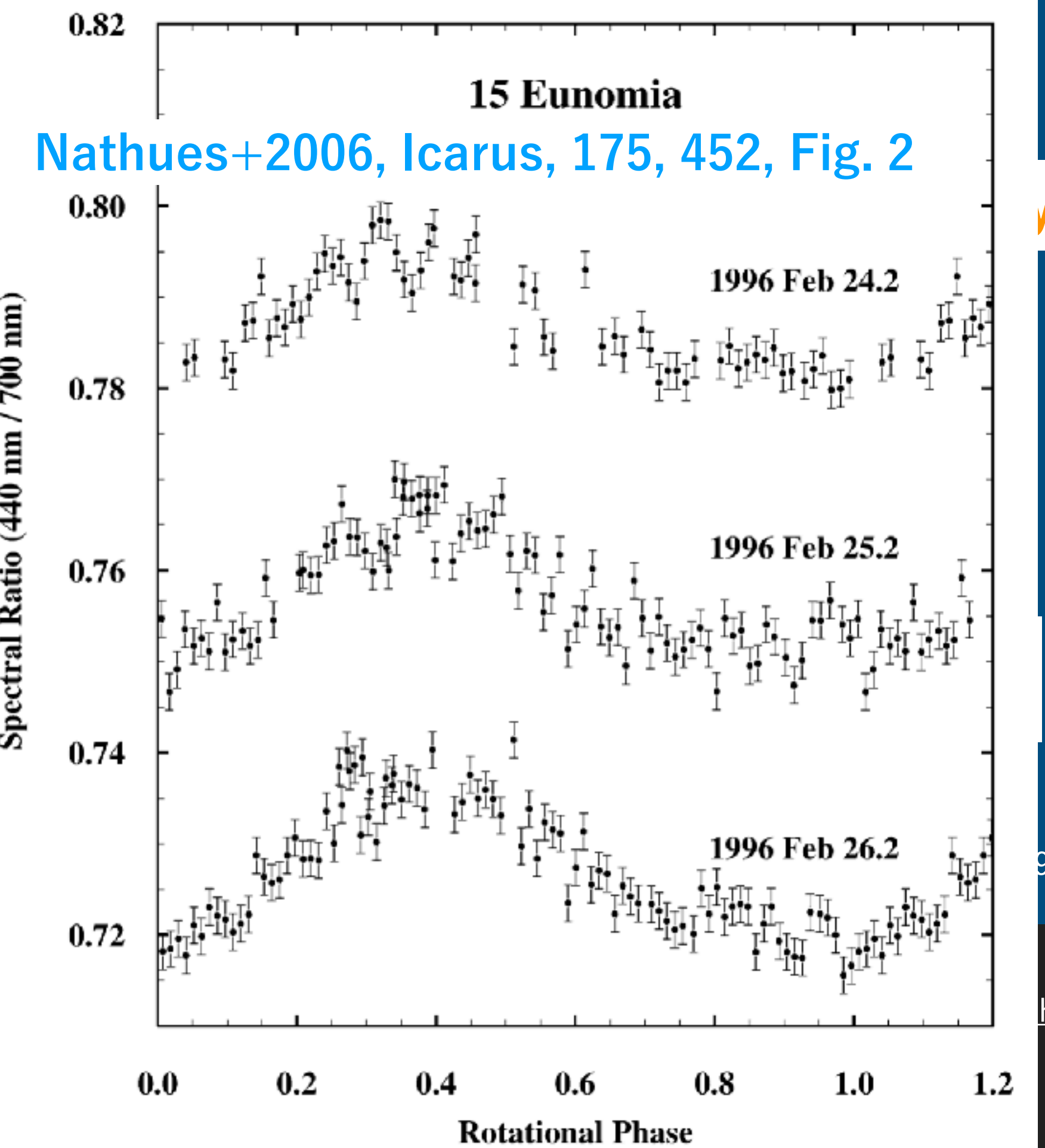
Difficulties and goals

- Small NEOs are hard to observe.
 - ▶ limited observational window (hours ~ a few days)
 - ▶ fast rotation (< 1 minute)
 - ▶ large apparent motion etc.

- The numbers of tiny NEOs are not enough for statistical study.

Goal of this study

- ① Estimate spectral types of tiny NEOs (tricolor + ToO).
- ② Search for spectral changes in rotationally resolved colors. (tricolor + ToO + time-resolved)



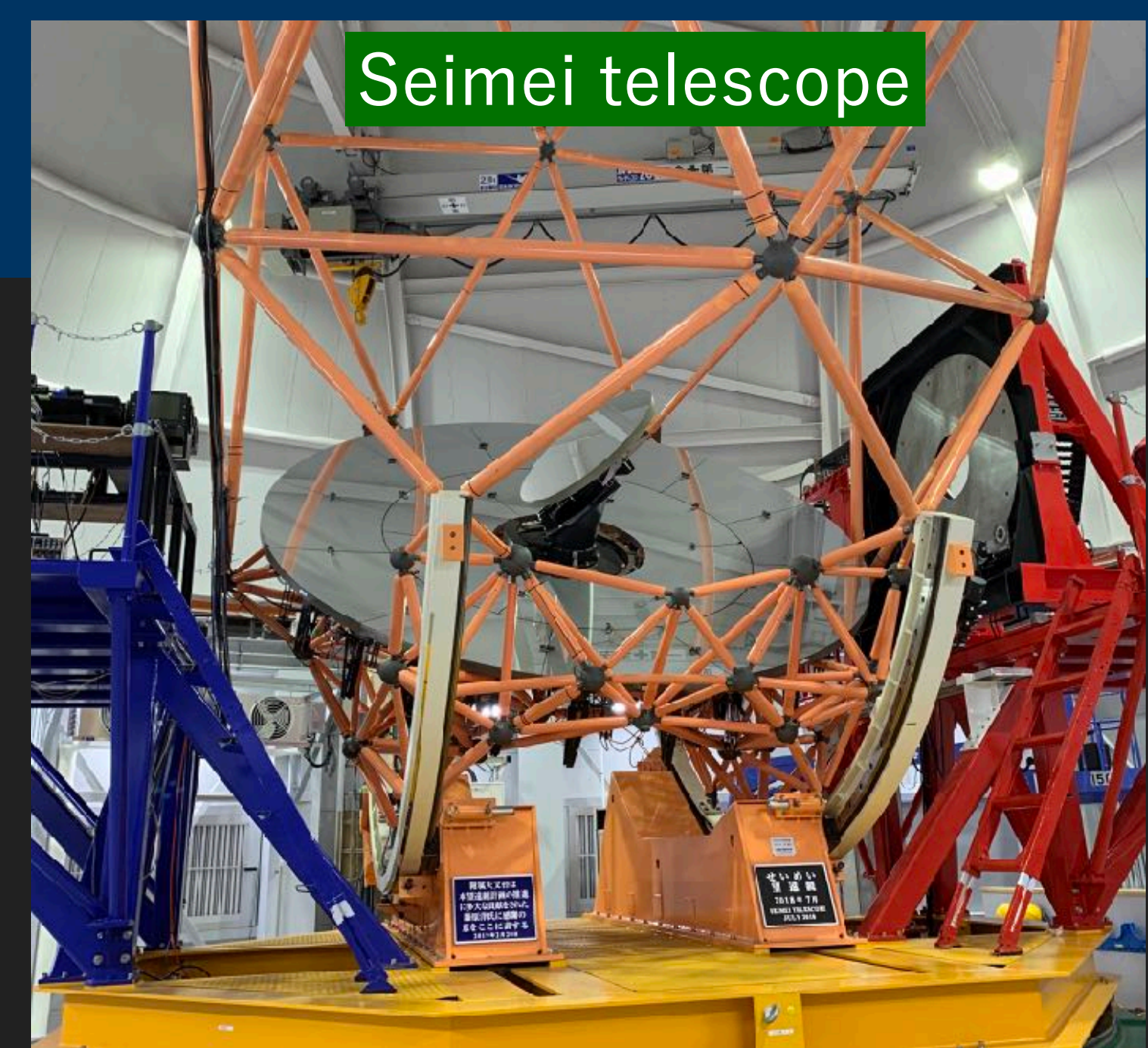
y time

html)

Observation

- Instrument : TriCCS

- ▶ TriColor CMOS Camera and Spectrograph
@ Kyoto University 3.8 m telescope (the Seimei telescope)
- ▶ g + r + i/z (Pan-STARRS) tricolor imager
- ▶ S/N=10, lim.mag ~ 18 @ g, r band, 1 sec exposure
- ▶ field of view 12.6 x 7.5 arcmin

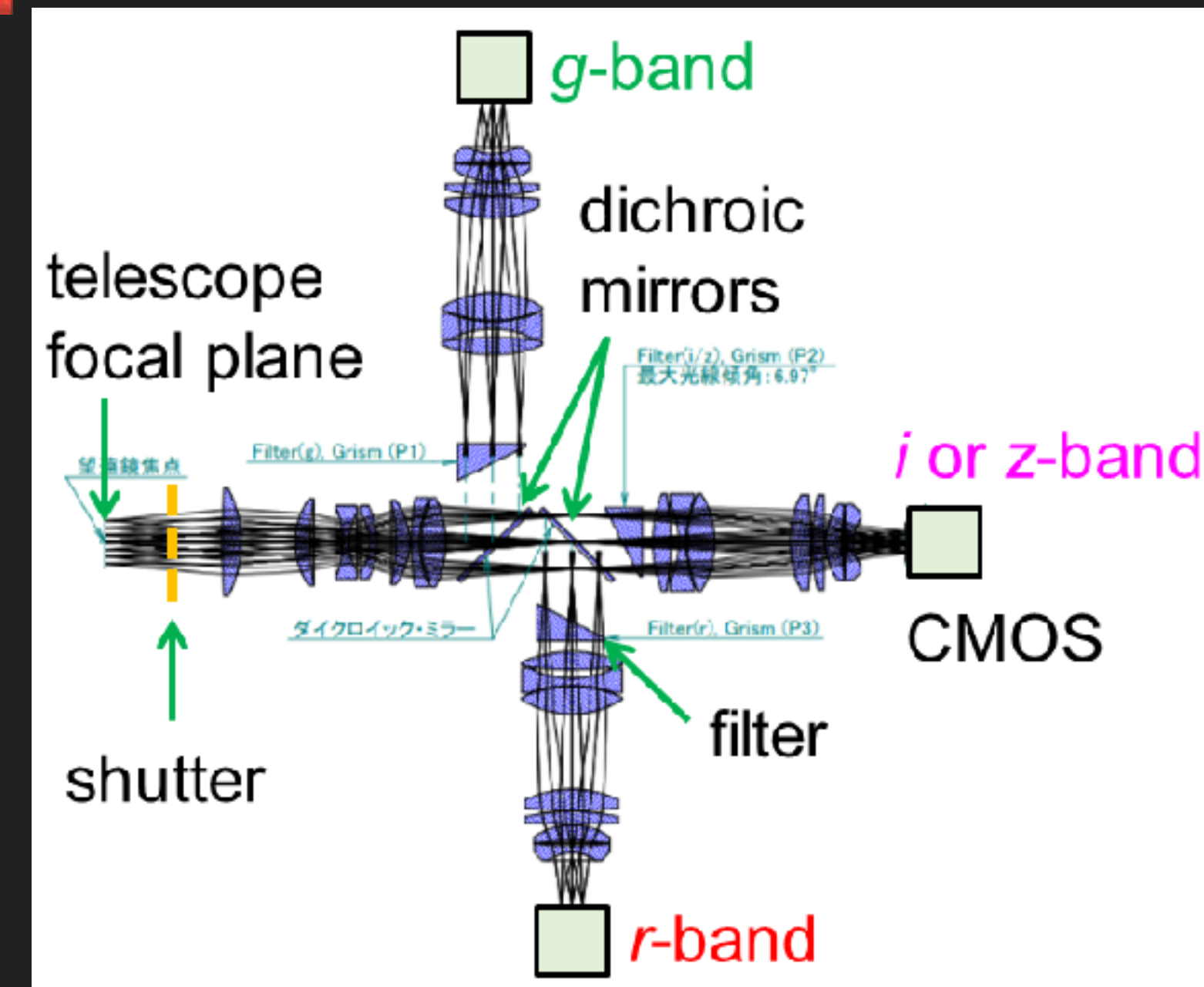


- Targets

- ▶ 5 NEOs (mean diameter ~ 50 m, soon after discovery ~ a few days)

- Observation Date & Time

- ▶ Feb 22-24th, Mar 8th, 2021
- ▶ 5 min – 60 min / NEO



Observed NEOs

Orbit diagram of a NEO (Feb 24th, 2021)

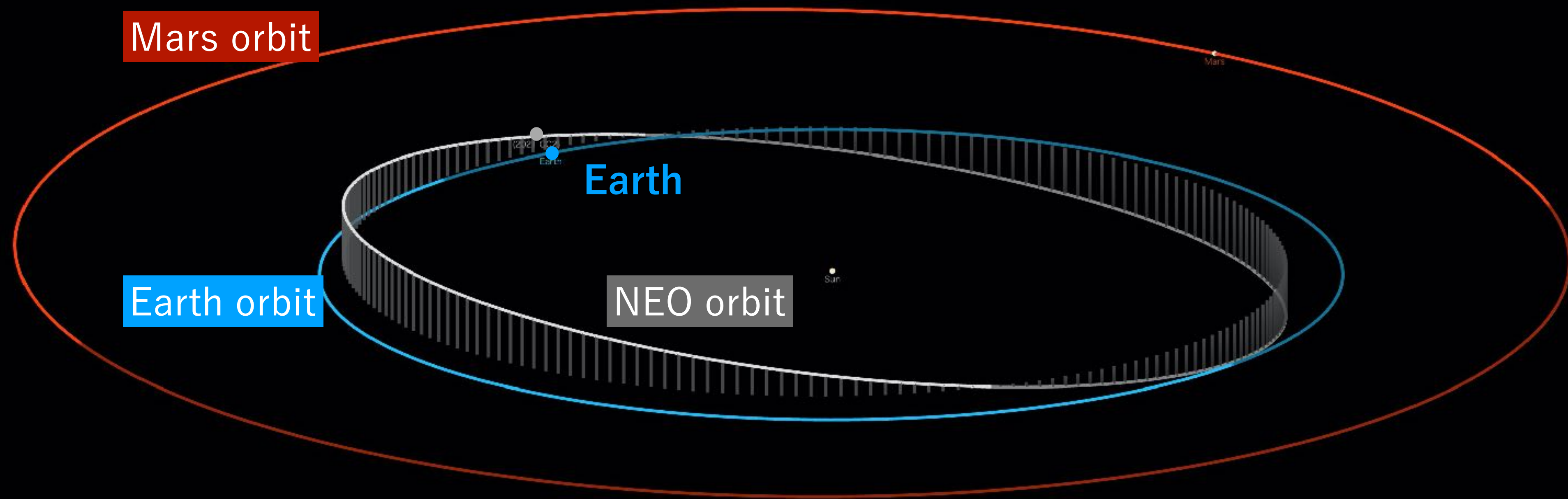
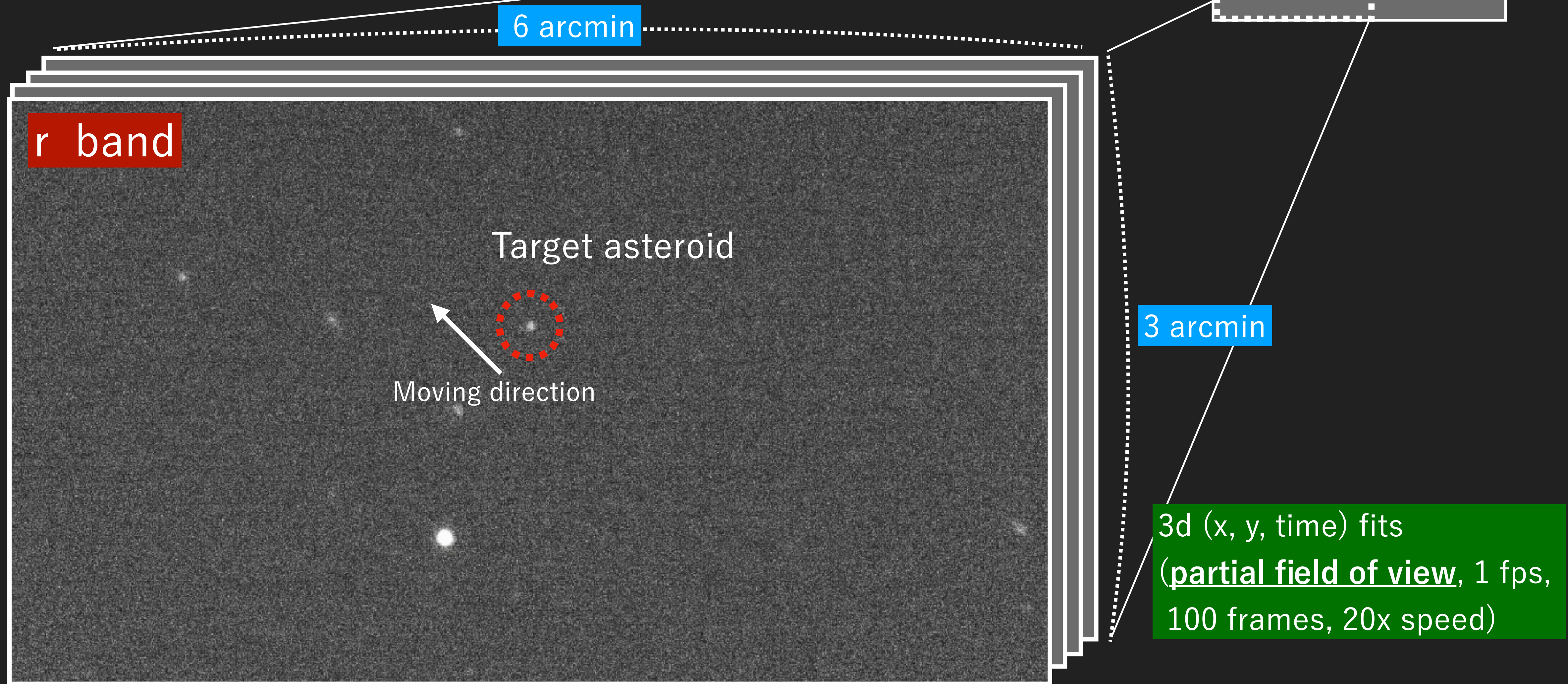


Image generated by JPL HORIZONS (<https://ssd.jpl.nasa.gov/horizons.cgi>)

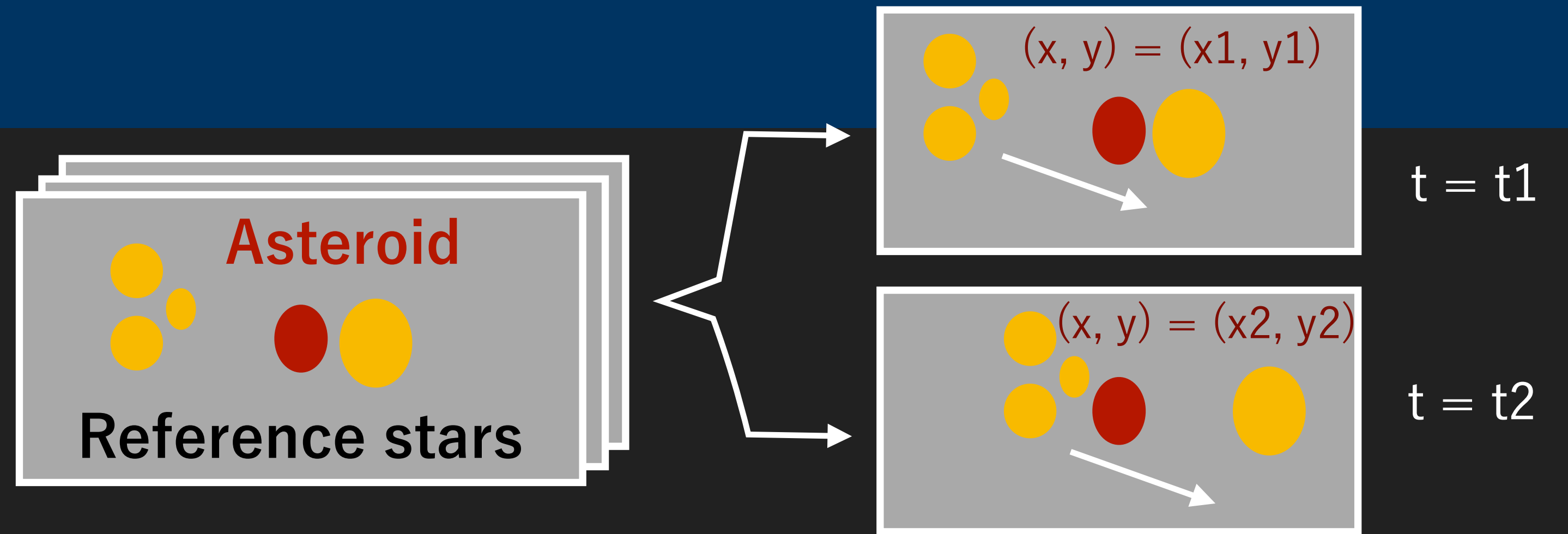
3D video cube

- Non-sidereal tracking



Data Reduction

- 0-1. bias & dark subtraction & flat-fielding
- 0-2. Split 3d (x, y, t) to 2d (x, y)
- 0-3. astrometry (WCS, astrometry.net*)



1. NEO location determination

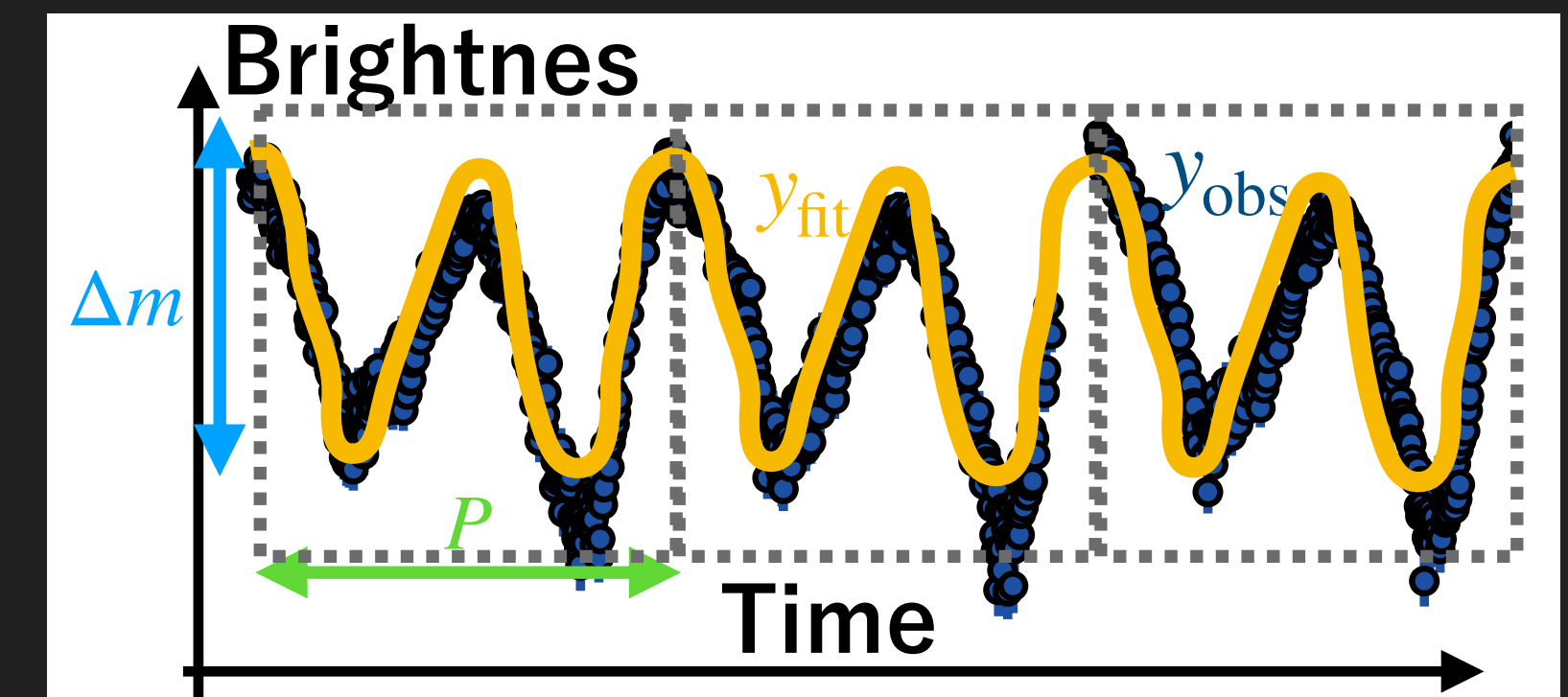
predict location by interpolation of
 $(x, y) = (x_1, y_1), (x_2, y_2) \dots$

2. Aperture photometry

- circle aperture (python, SEP**)
radius = 2 – 3 FWHM (~ 5 arcsec)
- relative photometry
with Pan-STARRS catalog***

3. Periodic analysis

- **Rotational periods** are estimated by Lomb-Scargle method
(Lomb 1976, Scargle1982, VanderPlas2018)



* Lang+2010, AJ, 131, 1682 ** Bertin+1996, AA, 117, 393, Barbary+2015, [10.5281/zenodo.15669](https://zenodo.org/record/15669)

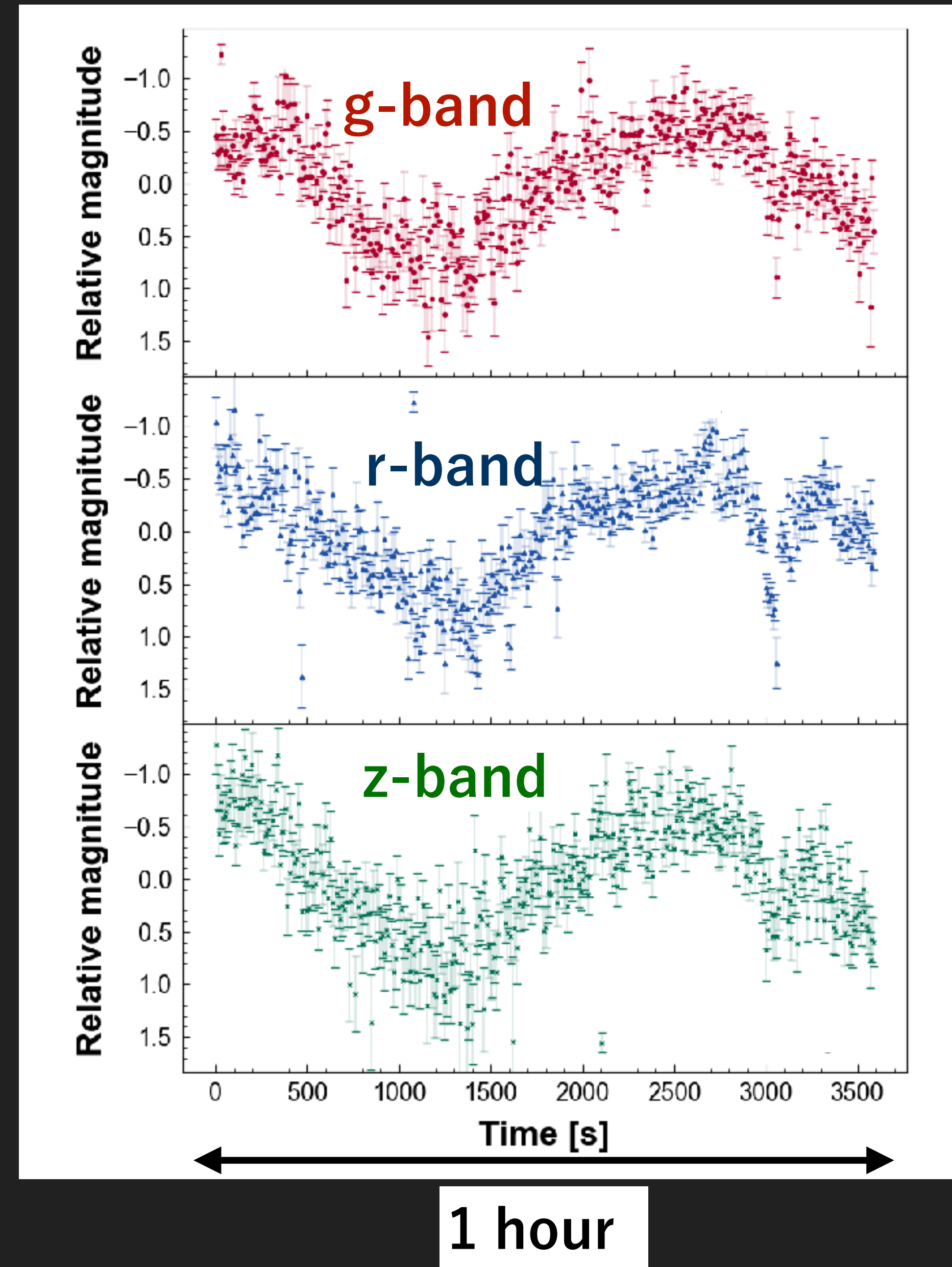
*** Chambers+2016, arXiv:1612.05560

Result 1. Light curve (NEO)

(Feb 24th, 2021)

t_{exp} : 10 sec (stack 1 sec x 10 frames)

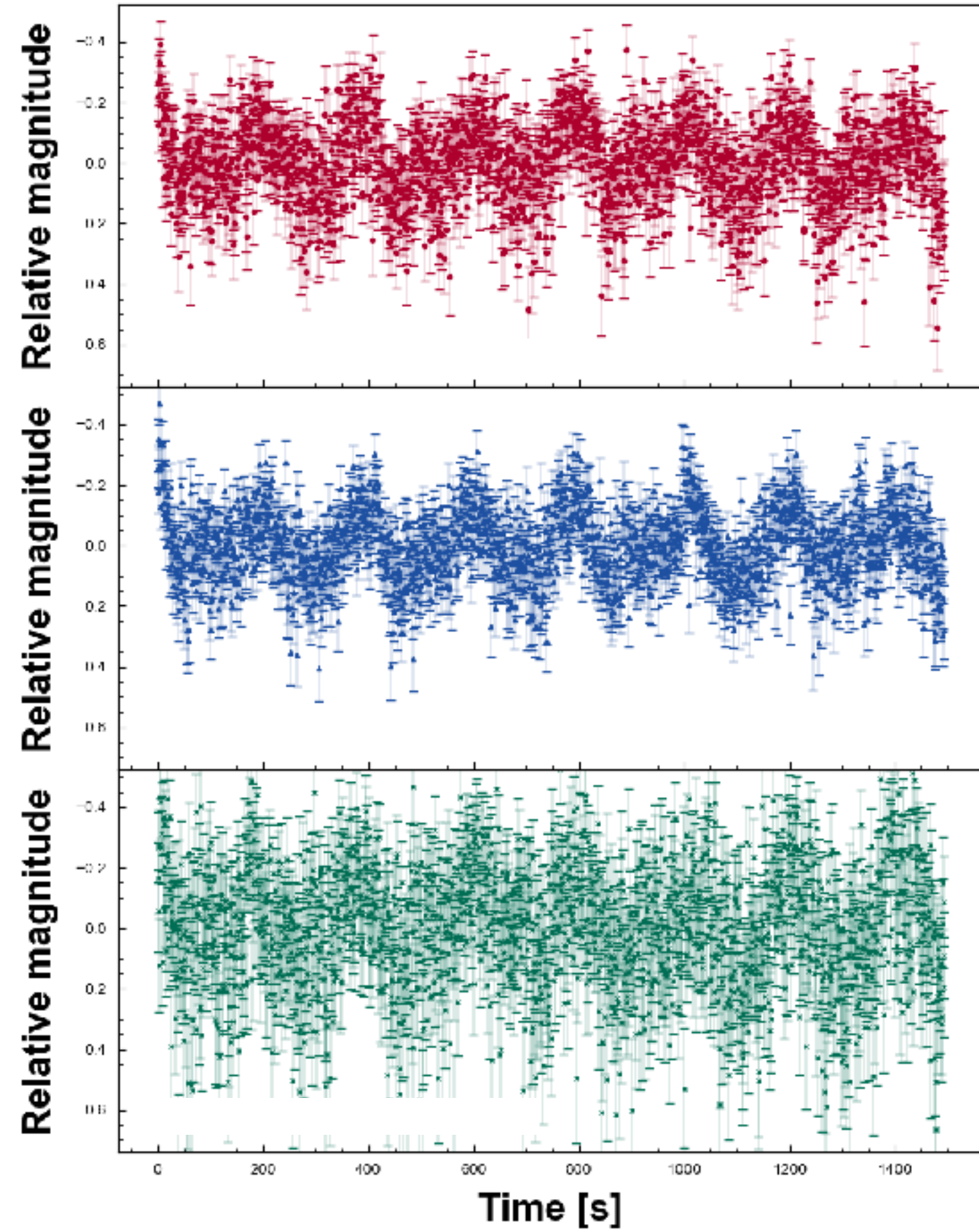
- The rotational period is not estimated in observing time (1 hour).
This is consistent with the period ~ 1.8 hour in a previous study*.



Result 1. Light curve (NEO)

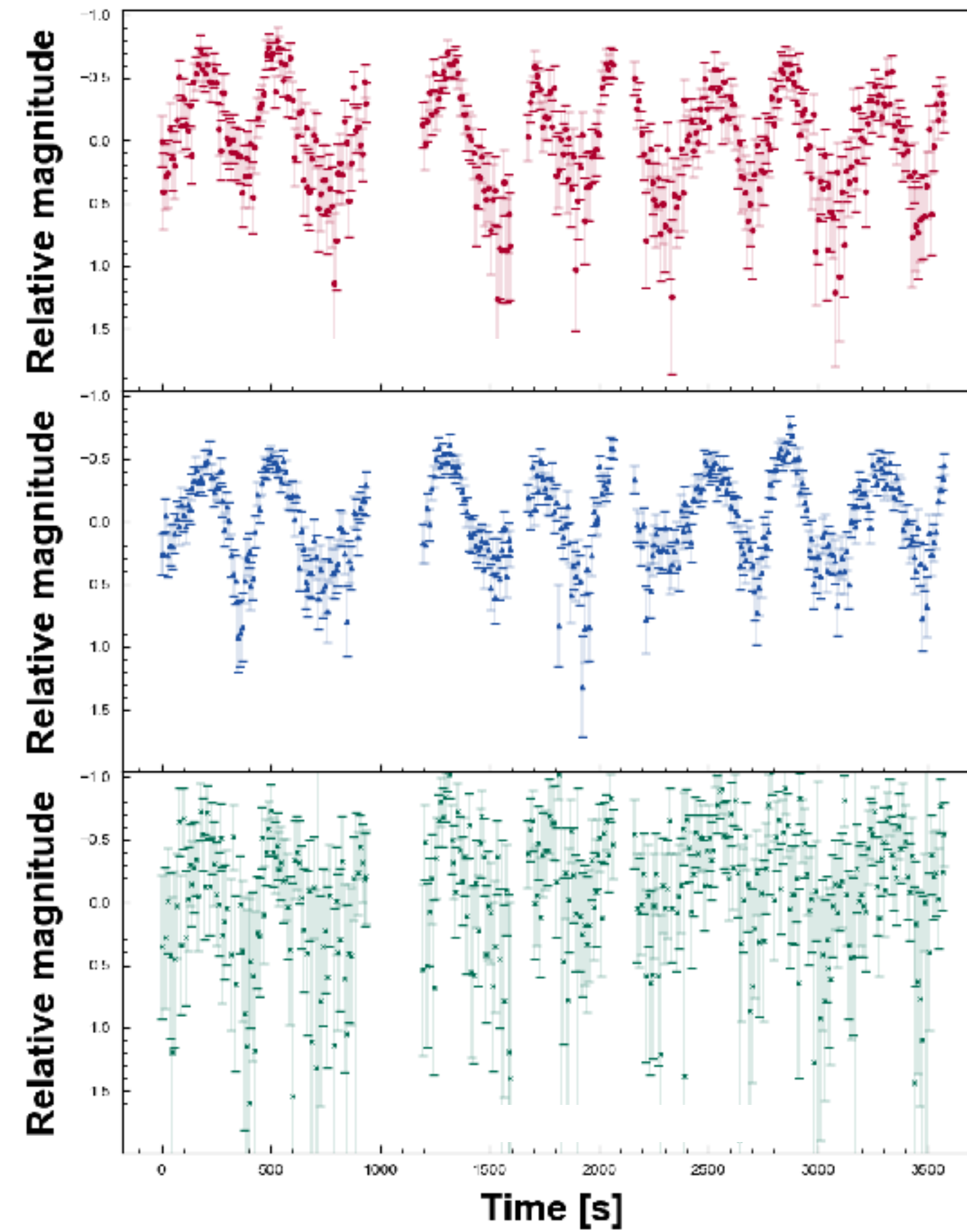
(Mar 8th, 2021)

$t_{\text{exp}} : 1 \text{ sec}$



(Feb 24th, 2021)

$t_{\text{exp}} : 10 \text{ sec (stack 1 sec x 10 frames)}$



Result 1. Light curve (Reference)

① 1 sec exp.

$g, r, i = (15.8, 15.4, 15.3)$
 $\sim 1.6\%, 1.2\%, 2.1\%$

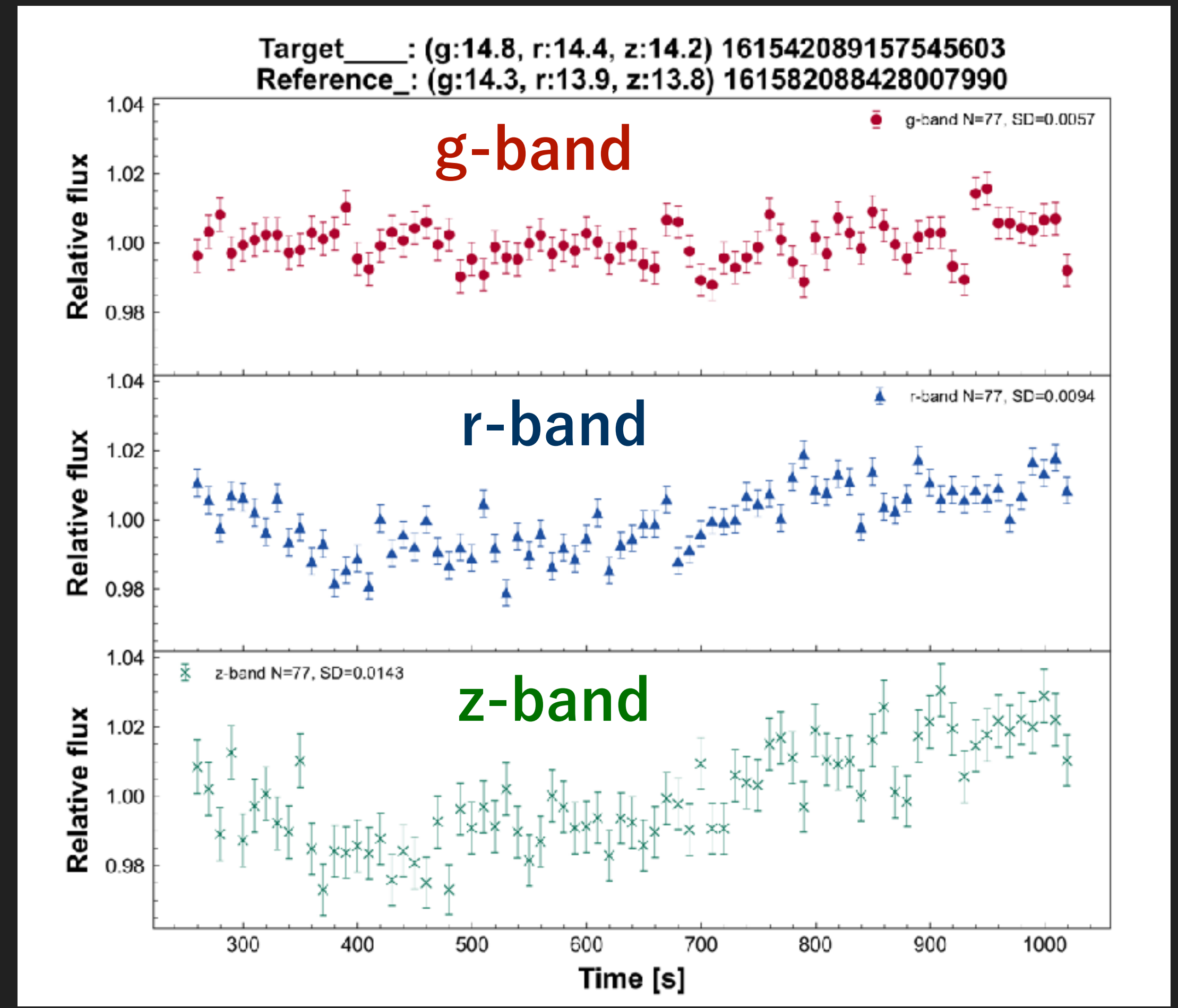
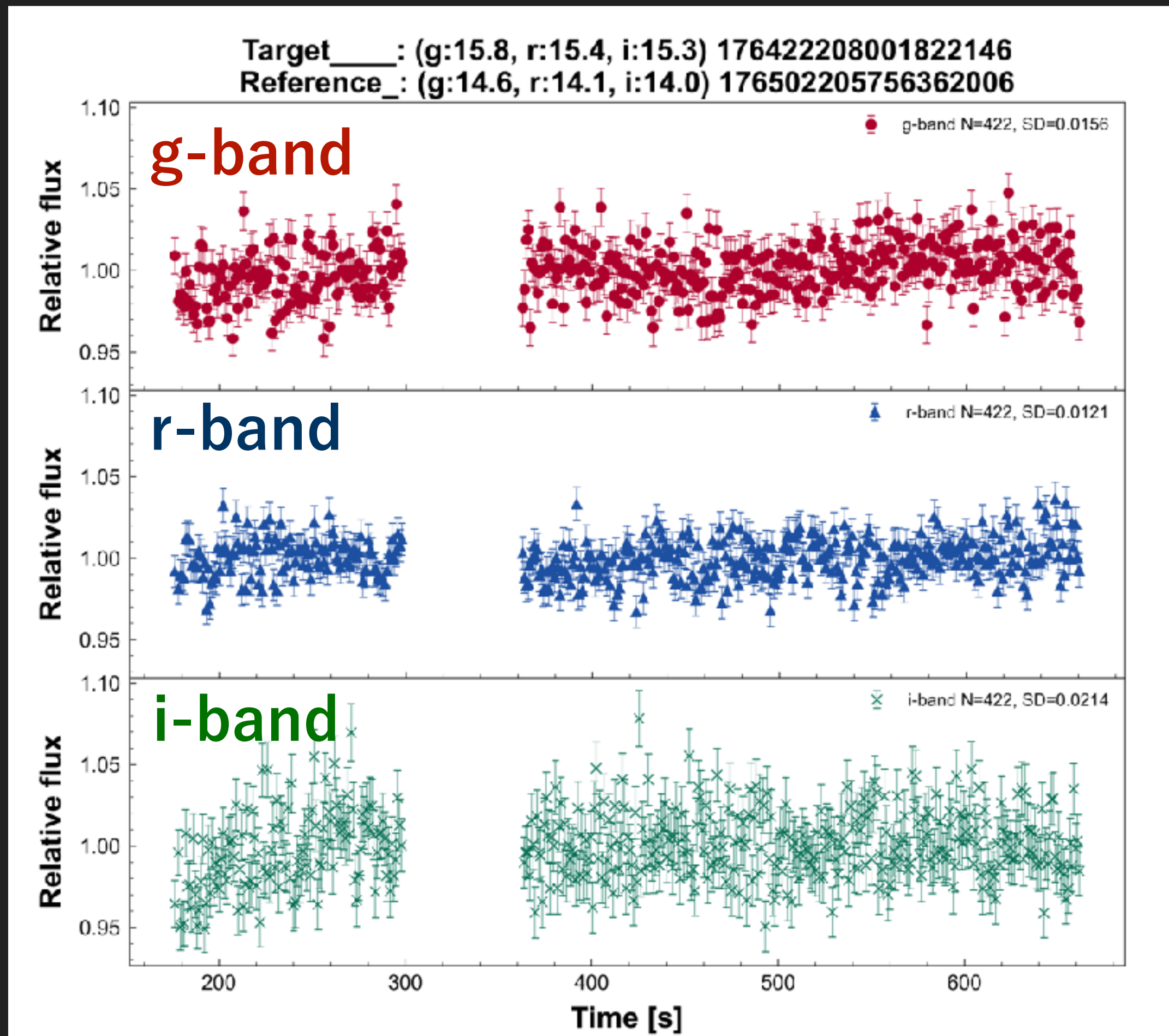
② 10 sec exp.

(stack 1 sec x 10 frames)

$g, r, z = (14.8, 14.4, 14.2)$
 $\sim 0.57\%, 0.94\%, 1.4\%$

normalized
 $\frac{F_{\text{target}}}{F_{\text{reference}}}$

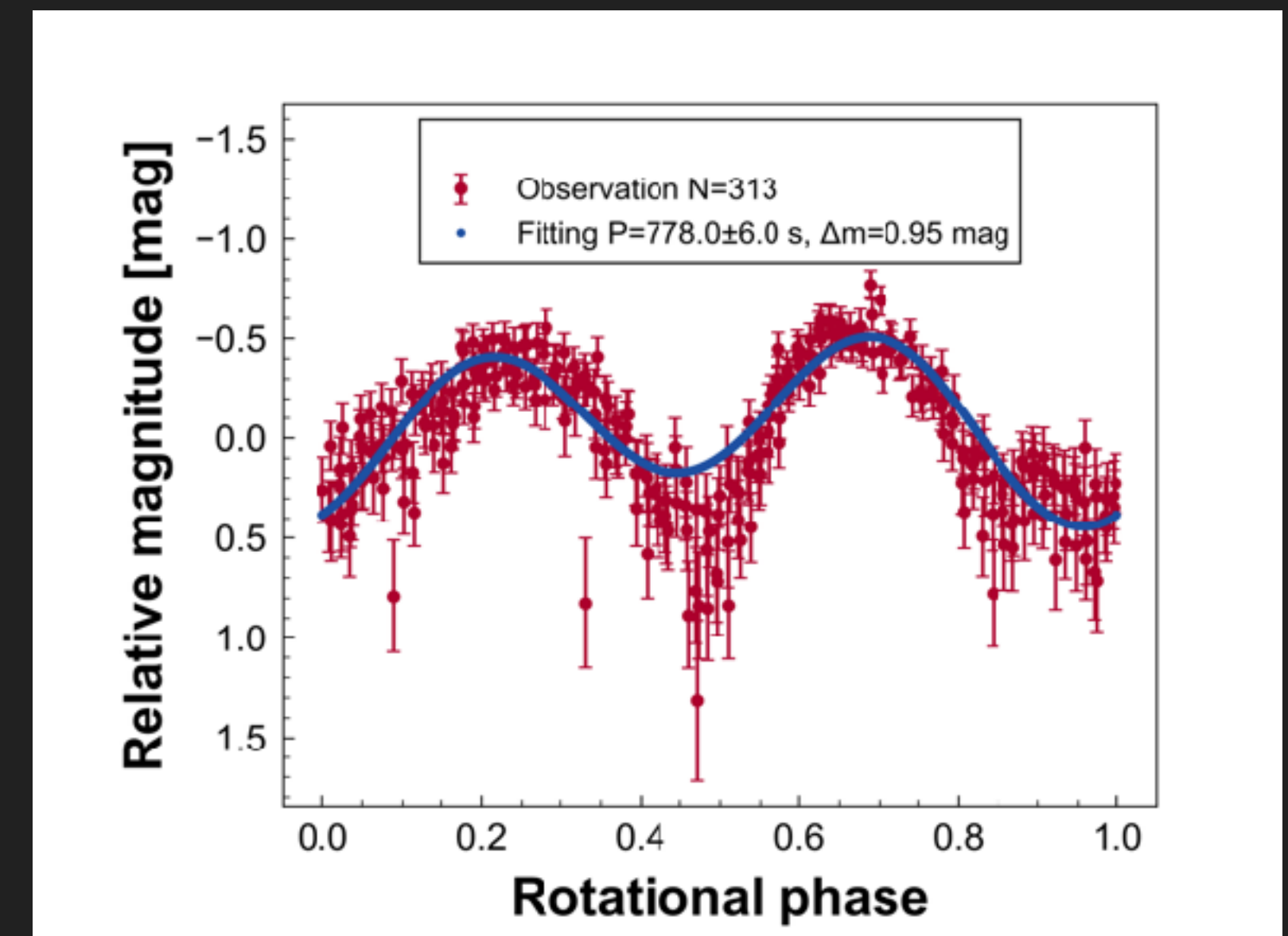
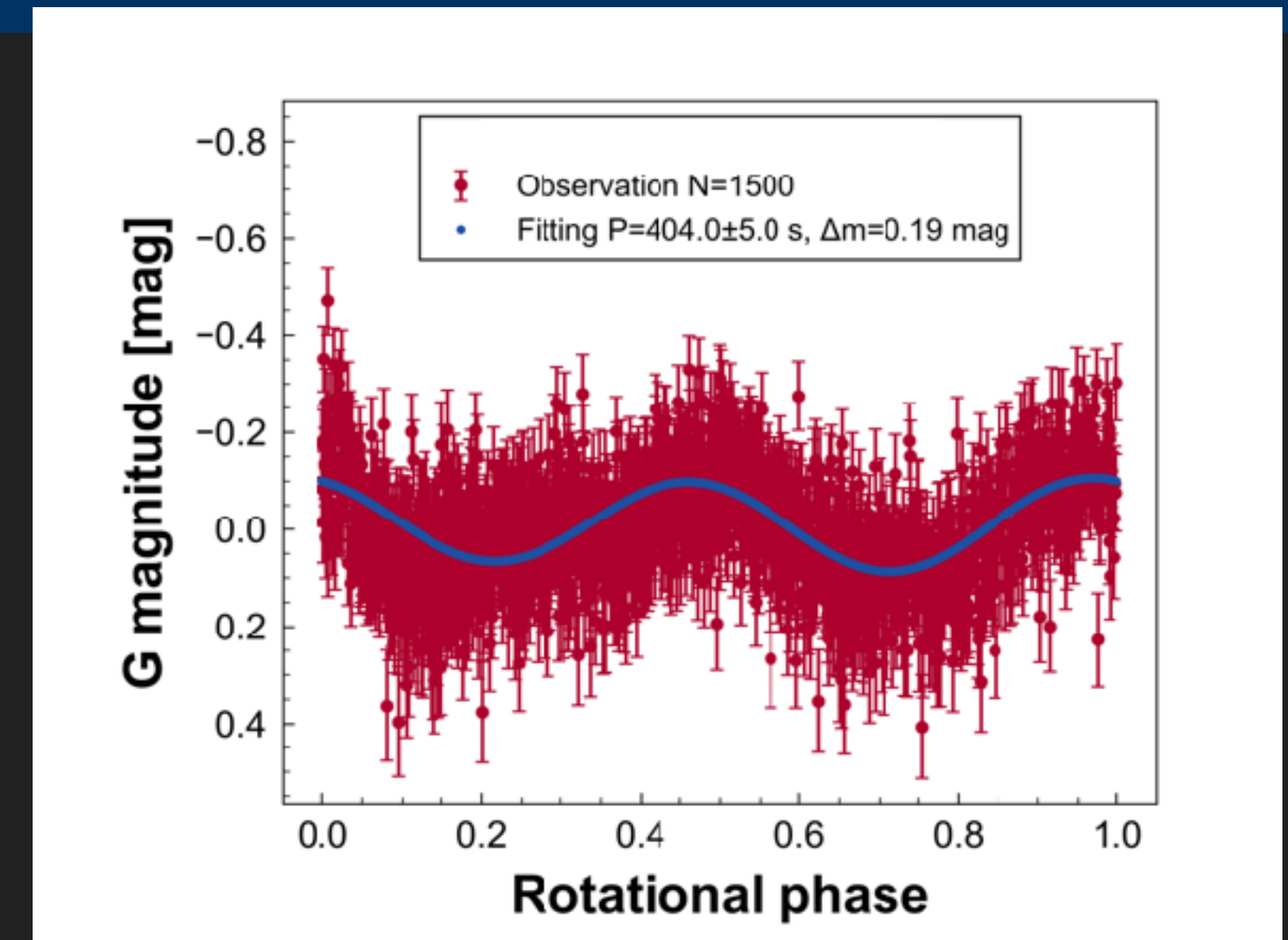
reference ●
target ●



Note: Ref. stars are moving at ~ 1 pix/sec due to the non-sidereal tracking.

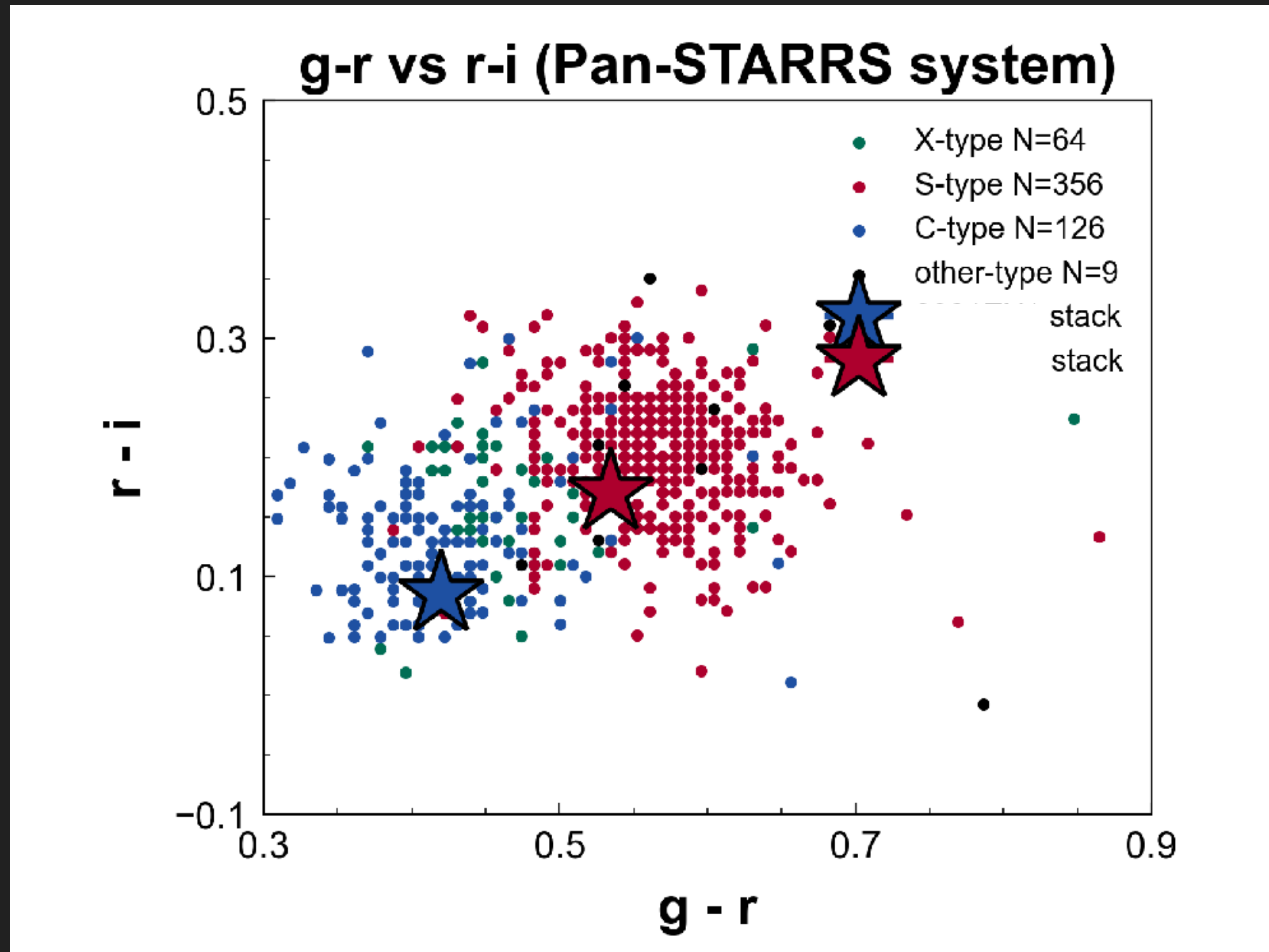
Result 2. Rotational period

- We derived the rotational periods of 2 NEOs :
 - ▶ (Mar 8th)
rot. $P = 404 \pm 5$ s
 - ▶ (Feb 24th)
rot. $P = 778 \pm 6$ s



Result 3. Spectral type

- (preliminary) Spectral types of 2 NEOs are estimated.

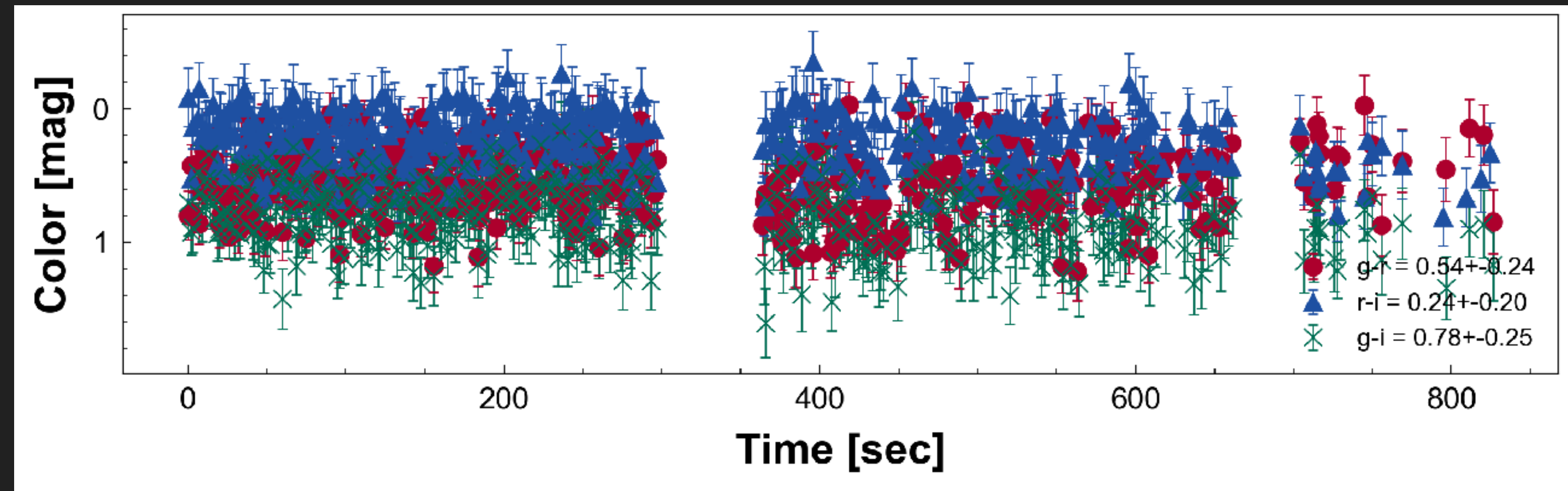


Reference objects are from SDSS Moving Object Catalog, Ivezic+2002, Proceedings of SPIE Vol. 4836, <http://faculty.washington.edu/ivezic/sdssmoc/sdssmoc.html>

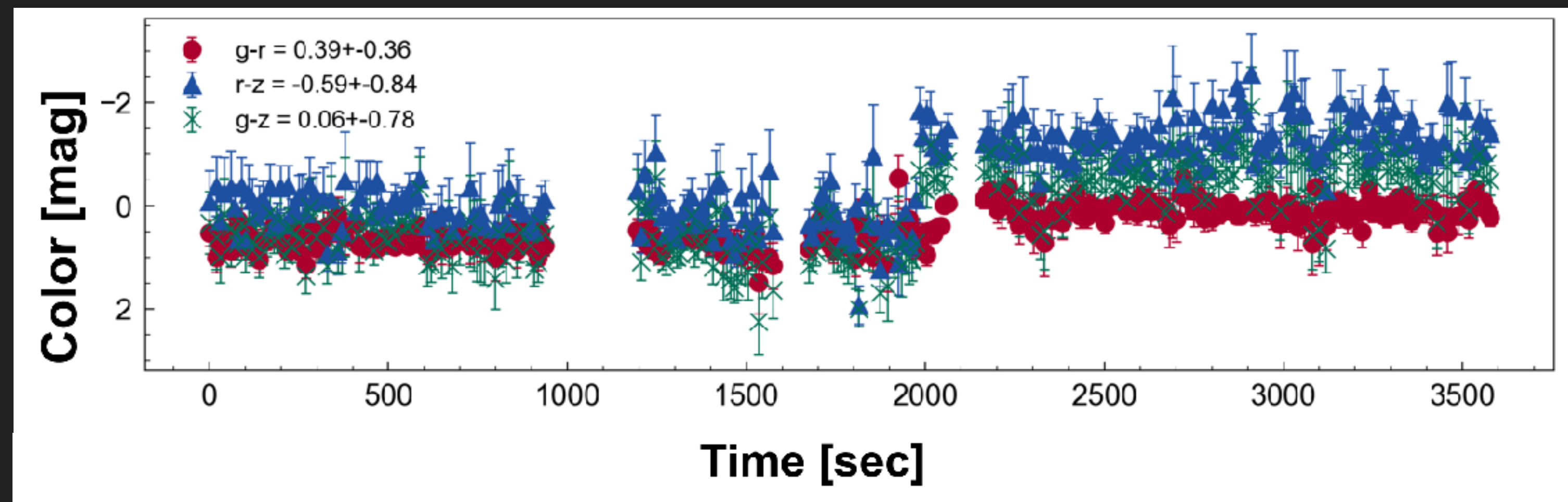
Result 4. Color variation

- (preliminary) No color variation is detected in all the 5 light curves.

(Feb 24th)



(Feb 24th)



More detailed
analysis are needed.
(c.f. color transformation)

Summary

- We conducted tricolor simultaneous photometric observations.
 - ▶ Instrument : Seimei telescope / TriCCS camera (g + r + i or z)
 - ▶ Mode : video observation at 1 fps
 - ▶ Target : 5 Near-Earth Objects (diameter < 100 m)
- We estimated both the rotational periods and the spectral types.
- No spectral change was found in all the 5 light curves (preliminary).
- Additional observations will be done in Semester 2021B.

2021B ToO観測へ向けて

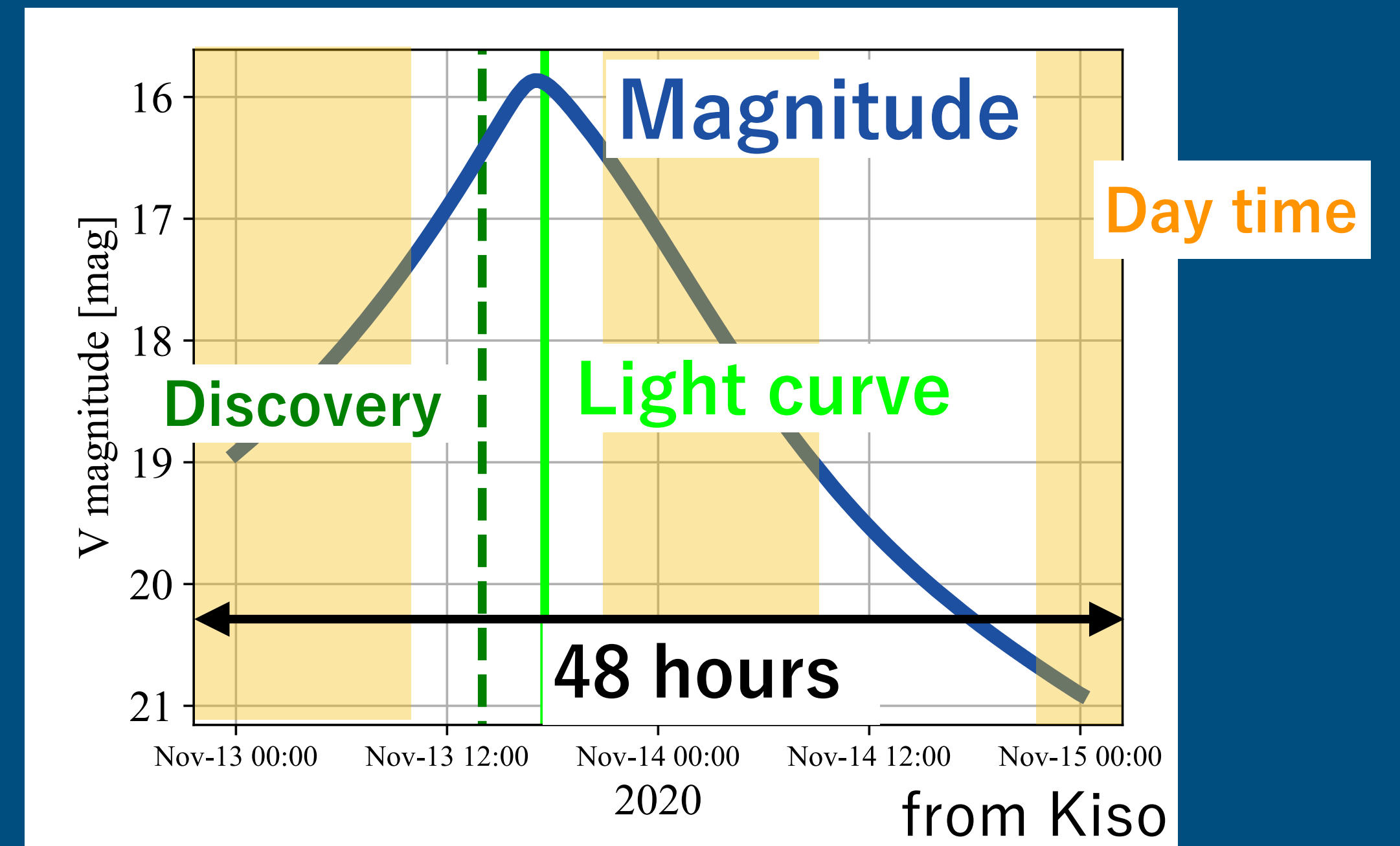
- ・ 当日または翌日のToO依頼が見込まれます。現地観測者や天文台の方にお願ひする際はよろしくお願ひします。
- ・ NEO観測のための手順書に従いコマンド入力をすると観測ができるように整備しました。(8/11に整備予定)

- ・ データ解析用scriptを公開しています。

1. TriCCS Data Reduction (tdr) https://bitbucket.org/jin_beniyama/triccs-data-reduction/src/master/
2. Moving Object Photometry (movphot) https://bitbucket.org/jin_beniyama/movphot/src/master/

【観測所への要望】

即日や翌日ToOの観測ですので、リモート観測を心待ちにしています。



Ephemeris of a tiny NEO (D = 5 m)