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On behalf of the Seimei-Kanata SN follow-up team & Tomo-e Gozen SN team



SN la diversity

- Standardized candles, but...
- Too much diversity to be a single population.
- Indications for different progenitors/explosions for different classes.
- Bulk standard + rare outliers?
 No!
- A large fraction of the nearby SNe have turned out to be peculiar beasts. Much more diverse than expected.



Frontiers in Transient Science Higher cadence



Cooke+ 2015

Unknown shorttime scale objects & just after the explosion. Larger samples **Unknown rare** types of explosions. Deeper Unknown faint objects & long term evolution.

See the poster by Morokuma et al.

New Time Domain Era



	Tomo-e SN Survey		
instrument	Tomo-e Gozen		
sensor	CMOS		
readout time	~0 sec		
period	2018/9-		
survey area [deg2]	10,000		
cadence	2 hours / 1 day		
exposure time / visit	3 sec		
depth	<mark>18 mag</mark> / 19 mag		
filter	no (~g+r)		
#(SBOs), #(SNe) / yr	5, 1000		
data storage	daily-stacked image SN cutout images		
reference	-		

Catch SNe in the first day. © Morokuma Pick up rare examples (with long-term observations).

The Seimei SN/Transient Program

Some numbers		© Tanaka		* Numbe		
Distance (Mpc)	Volume (Mpc³)	# of galaxies	# of SNe (yr-1)	mag (abs mag -15 mag)	mag (abs mag -13 mag)	
10	4 x 10 ³	40	0.4	15.0	17.0	
30	1 x 10 ⁵	10 ³	10	17.4	19.4	~ 5 / yr.
50	5 x 10⁵	5 x 10 ³	50	18.5	20.5	
70	1 x 10 ⁶	1 x 104	100	19.2	21.2	~ 50 / y
100	4 x 10 ⁶	4 x 10 ⁴	400	20.0	22.0	
200	3 x 10 ⁷	3 x 10 ⁵	3,000	21.5	23.5	
500	5 x 10 ⁸	5 x 10 ⁶	50,000	23.5	25.5	

Roughly, ~ 10 classical nights + 10 ToO nights / semester.
soon-after-the-explosion SNe up to ~ 30 Mpc (golden set).
well-characterized SNe up to ~ 70 Mpc (a hunt for outliers).
Kawabata-san, Yamanaka-san, Isogai-san, ...
So far mostly fueled by public surveys; Tomo-e to come.

Data taken also within the OISTER Rapidly evolving + high velocity (HV) SN Ia 2019ein



t= -6 davs

1.6

1.8

10000

0.8

1.2

 $\Delta m_{15}(B) \text{ (mag)}$

No rapidly-evolving HV SNe well observed in the rising phase before. "highest velocity" for SN Ia ever observed.

Implications for the SN Ia diversity and the explosion



Diversity:

HV SNe are not a one-parameter family; not only the velocity but also the speed of the evolution does matter. The explosion model: Extended nuclear burning out to the surface region in HV SNe.





- In the presentation, properties of two SNe Ia were discussed, based on the Seimei/Kanata/Tomo-e data.
- The contents are in preparation, one by Kawabata et al. and the other by Kawabata et al. and Jiang et al.
- These are removed from the online proceedings here.

Future (Spectroscopy) Perspectives

Some numbers		© Tanaka		* Numbers for all sky				
Distance (Mpc)	Volume (Mpc³)	# of galaxies	# of SNe (yr-1)	mag (abs mag -15 mag)	mag (abs mag -13 mag)	NOW F	ull spe Kool <u>s</u>	€C
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Infant sample: soon-after-the-explosion SNe (down to -15 mag).
Basic sample: good characterization (a hunt for outliers: -17 mag).
We are now able to find weirdos at a rate of 1 in 50 SNe.
Will be able to do it for 1 / 200 SNe with full-spec Seimei + Kools.
Further down to 1 / 1,000 with TriCCS to come.
Need a number of nights + human resource for science.

See the talk by Ohta et al.

TriCCS: new imager + spectrograph



3 channel CMOS camera (fully funded, 2021~) Imager→spectrograph (fully funded, 2023~?) A new postdoc, ~ Apr 2021? TriCCS pipeline + obs (Job ad @ ~ Dec 2020) Spectra @ 1 day for ~ 30 SNe / yr within 60Mpc (currently <10 SNe reported)



Take Away

- Peculiar beasts beyond known diversities have been regularly and routinely found.
- Infant phase observations routinely performed with Seimei and Kanata telescopes.
- New insight obtained:
 - High velocity SNe: not a single param. family; extended burning.
 - Two other SNe Ia: expanding the diversity; insights into the progenitor systems and the explosion mechanisms.
- Will produce **a complete census** with coming update in Seimei and instrument. **Stay tuned!**